

Real-Time System

- CfA
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 - Randall Wayth – imaging & calibration
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- IIC (Harvard)
 - Richard Edgar – GPU coding
- Melbourne
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Outline

- Overview and Status

- Preparation for a March '09 deployment

- Work to be done

- GPU conversion
- Algorithm development
- Timeline
- Risk

- Strategy for build-out

- Porting to a larger cluster
- Algorithm development
- Risk

Status of 32T RTS

- All required hardware is in WA.
- Basic pipeline regularly processes MAPS data (CPU version).
- Much of the C-code has been converted to the NVIDIA Compute Unified Device Architecture, CUDA.
- RTS distribution runs in either CPU or CPU+GPU mode (not a full CUDA pipeline yet).
- CPU version has been used to calibrate and image visibilities from the November 2008 site visit. Several hacks need to be worked into the system.

32T RTS Tasks

- Interface between the Correlator and the RTS (Stewart + CfA).
- Head node \Leftrightarrow compute nodes code (Stewart + CfA).
- Complete CUDA pipeline:
 - Gridder and imager both implemented (gridder needs optimisation).
 - Calibrator Measurement Loop implemented (needs optimisation and testing).
 - Stokes conversion implemented (needs optimisation and testing).
 - Use simple & fast regridding (bilinear – possibly on the CPU).
- Implement fixes for real data:
 - RFI detection / flagging (currently setting `inv_var` weights to zero).
 - Direction-independent calibration using peel solutions of a single bright source.
 - Variable calibrators, resolved calibrators, ...
- Much testing once real data is flowing!

32T RTS Timeline

- January:
 - Finish master node.
 - Finish conversion to CUDA (expect speed up factor of 5-10).
- February:
 - Attach to test-bench at Haystack (integration of UMelb & CfA code).
 - Continue to optimise routines.
 - Continue to test and develop real-time calibration.
- March:
 - Install code on the machines in Boolardy.
 - Test, fix, test, fix, ...

32T RTS Risk \approx GPUs

- I/O Concerns:
 - INPUT ✓ OUTPUT ✓ (sustained transfers ?)
 - CALIBRATION ? ✓
- Memory Concerns ✓
- Processing Concerns:
 - 10 calibrators \rightarrow $< 65 \text{ Gflops}/8\text{sec}/\text{GPU} = 57 \text{ Gflops}/\text{sec}/\text{GPU}$ (7 cycles/flop). ✓
 - Mostly Stokes conversion (fewer baselines than 512T but similar image size).
 - Much of the Stokes conversion code can be pre-computed every few minutes.
- Ionospheric Faraday rotation ?
- Primary beam correction ?

Tasks (build out from 32T)

- Continue to optimise algorithms.
- Primary beam models
 - Stability of beam shapes.
 - Time and frequency derivatives.
 - Parameterisation.
- Ionospheric models
 - Smallest and largest scales: temporal / angular / refractive / ...
- Incorporate Justin Kasper's ionospheric FR technique.
- Investigate various gridding kernels for different science objectives (best case / worst case effect on image quality, deconvolution complexity, ...)

Risks 1 (build out from 32T)

- GPU performance less than anticipated
 - Will learn from 32T.
 - Reduce number of calibrators, reduce baselines used in the fits, ...
 - Simplify regridding / Stokes conversion.
- Not enough calibrators
 - **Limited FLOPS**: increase FLOPS.
 - **Limited by confusion noise**: use a more complicated model for the pre-peel step.
 - **Small scale structure in the ionosphere**: use sources from the images.
 - Key science applications require 100s of sources subtracted in real-time.

Risks 2 (build out from 32T)

- Variable calibrators
 - **Spectral variations:** 5-10% errors from catalogues, or determine during absolute calibration process.
 - **Temporal variations:** model variation, or don't use in primary beam fits.
 - **Ionospheric scintillation:** detect and don't use in primary beam fits.
- Other real-time applications
 - Need to manage real-time resources.